

Einstein's velocity addition

("Zur Elektrodynamik bewegter Körper", Ann. Phys. 17 (1905), 891-921):

@ p. 906:

$$U = \frac{\sqrt{v^2 + w^2 + 2vw \cos \alpha - \left(\frac{vw \sin \alpha}{c}\right)^2}}{1 + \frac{vw \cos \alpha}{c^2}}$$

dimensionless:

$$\frac{U}{c} = \frac{\sqrt{\frac{v^2}{c^2} + \frac{w^2}{c^2} + 2\frac{v}{c}\frac{w}{c} \cos \alpha - \left(\frac{v}{c}\frac{w}{c} \sin \alpha\right)^2}}{1 + \frac{v}{c}\frac{w}{c} \cos \alpha}$$

which is:

$$\beta_0 = \frac{\sqrt{\beta_1^2 + \beta_2^2 + 2\beta_1\beta_2 \cos \alpha - \beta_1^2\beta_2^2 \sin^2 \alpha}}{1 + \beta_1\beta_2 \cos \alpha}$$

Note:

$\{\alpha, \beta_0, \beta_1\}$ observed in stationary frame,
 β_2 in frame moving at β_1 in stat. frame.

Parallel:

$$(\alpha = 0) \Rightarrow \beta_0 = \frac{\sqrt{\beta_1^2 + \beta_2^2 + 2\beta_1\beta_2}}{1 + \beta_1\beta_2} = \frac{\beta_1 + \beta_2}{1 + \beta_1\beta_2}$$

Perpendicular:

$$\left(\alpha = \frac{\pi}{2}\right) \Rightarrow \beta_0 = \sqrt{\beta_1^2 + \beta_2^2 - \beta_1^2\beta_2^2}$$

Lorentz factor:

$$\gamma = \frac{1}{\sqrt{1-\beta^2}} \therefore \gamma^2(1-\beta^2) = 1 \therefore \gamma^2 - \gamma^2\beta^2 = 1 \therefore \gamma^2 - 1 = \gamma^2\beta^2$$

$$\therefore \beta^2 = \frac{\gamma^2 - 1}{\gamma^2} = 1 - \frac{1}{\gamma^2}$$

Perpendicular velocity addition:

$$\beta_0 = \sqrt{\beta_1^2 + \beta_2^2 - \beta_1^2\beta_2^2}$$

$$\gamma_0 = \frac{1}{\sqrt{1-\beta_0^2}} = \frac{1}{\sqrt{1-\beta_1^2-\beta_2^2+\beta_1^2\beta_2^2}} = \frac{1}{\sqrt{1-\left(1-\frac{1}{\gamma_1^2}\right)-\left(1-\frac{1}{\gamma_2^2}\right)+\left(1-\frac{1}{\gamma_1^2}\right)\left(1-\frac{1}{\gamma_2^2}\right)}}$$

$$= \frac{1}{\sqrt{1-1+\frac{1}{\gamma_1^2}-1+\frac{1}{\gamma_2^2}+1-\frac{1}{\gamma_1^2}-\frac{1}{\gamma_2^2}+\frac{1}{\gamma_1^2}\cdot\frac{1}{\gamma_2^2}}} = \frac{1}{\sqrt{\frac{1}{\gamma_1^2\gamma_2^2}}} = \gamma_1\gamma_2$$

Parallel:
$$\beta_0 = \frac{\beta_1 + \beta_2}{1 + \beta_1 \beta_2}$$

perpendicular:
$$\gamma_0 = \gamma_1 \gamma_2$$

suffix 0: object's velocity w.r.t. stationary observer;
 suffix 2: object's velocity w.r.t. some moving ref. point P ;
 suffix 1: P 's velocity w.r.t. stationary observer
 (consistent with the aforementioned, but 1 & 2 are of course interchangeable).

$$(\beta \rightarrow 0) \Leftrightarrow (\gamma \rightarrow 1) \quad (\beta \rightarrow 1) \Leftrightarrow (\gamma \rightarrow \infty)$$

If $\{\beta_1, \gamma_1\} \rightarrow \{1, \infty\}$ **The more an object's velocity w.r.t. P approaches c ,**
 then $\{\beta_0, \gamma_0\} \rightarrow \{1, \infty\}$ **the more its velocity w.r.t. stat. obs. approaches c ,**
 totally indep. of $\{\beta_2, \gamma_2\}$ **i.e. the more its velocity becomes absolute, namely c .**

Relativity yields absoluteness...

Lorentz contraction of a moving object:

l := length of moving object, measured in direction of motion;

$$l_{\text{seenByStationary}} = l_{\text{seenByComoving}} \sqrt{1 - \beta^2}$$

If *you* are speeding, then from *your* perspective the street is moving, hence contracted by the same factor $\sqrt{1 - \beta^2}$.

The faster you go, the shorter the street appears to you
(cf. light experiencing zero time and distance).

Ever realised this street extends to "EDGE" of COSMOS?

**Would you reach the speed of light,
then you would — *at this very speed of light* —
*instantly smack against the "edge" of the universe!***

(Whatever the latter might be).